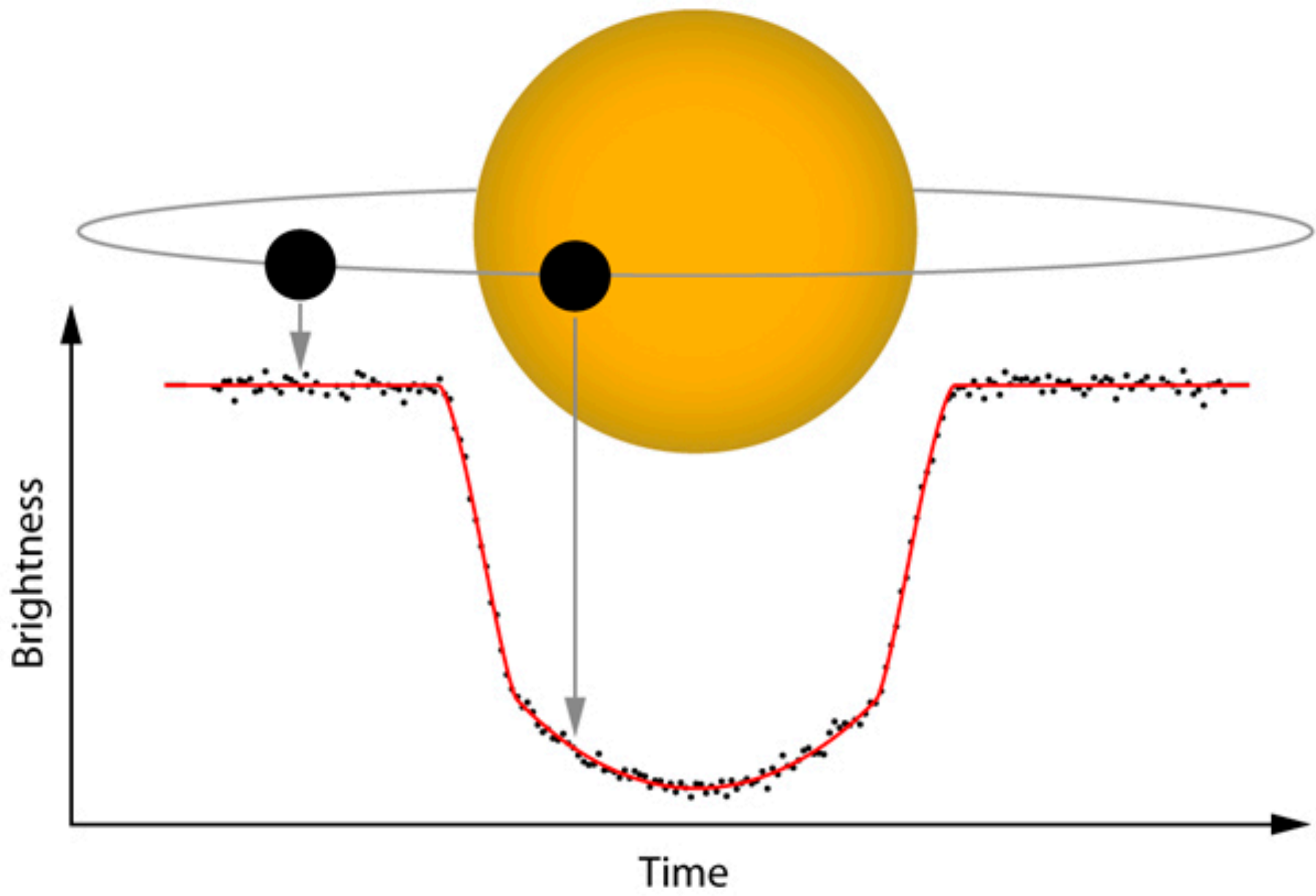


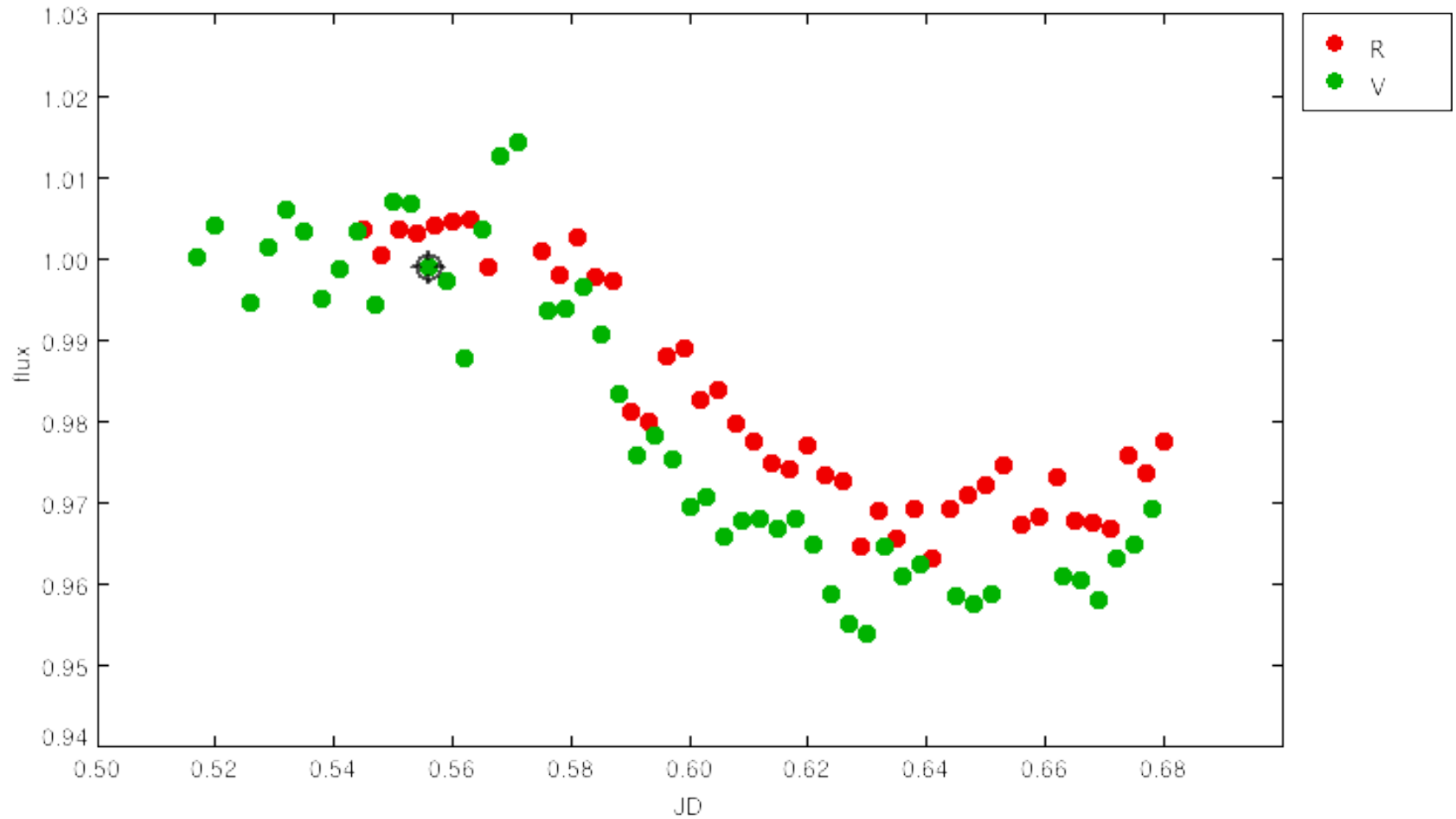
Multicolour photometry of transiting exoplanets

Juan Fabregat

GEOS meeting 2017



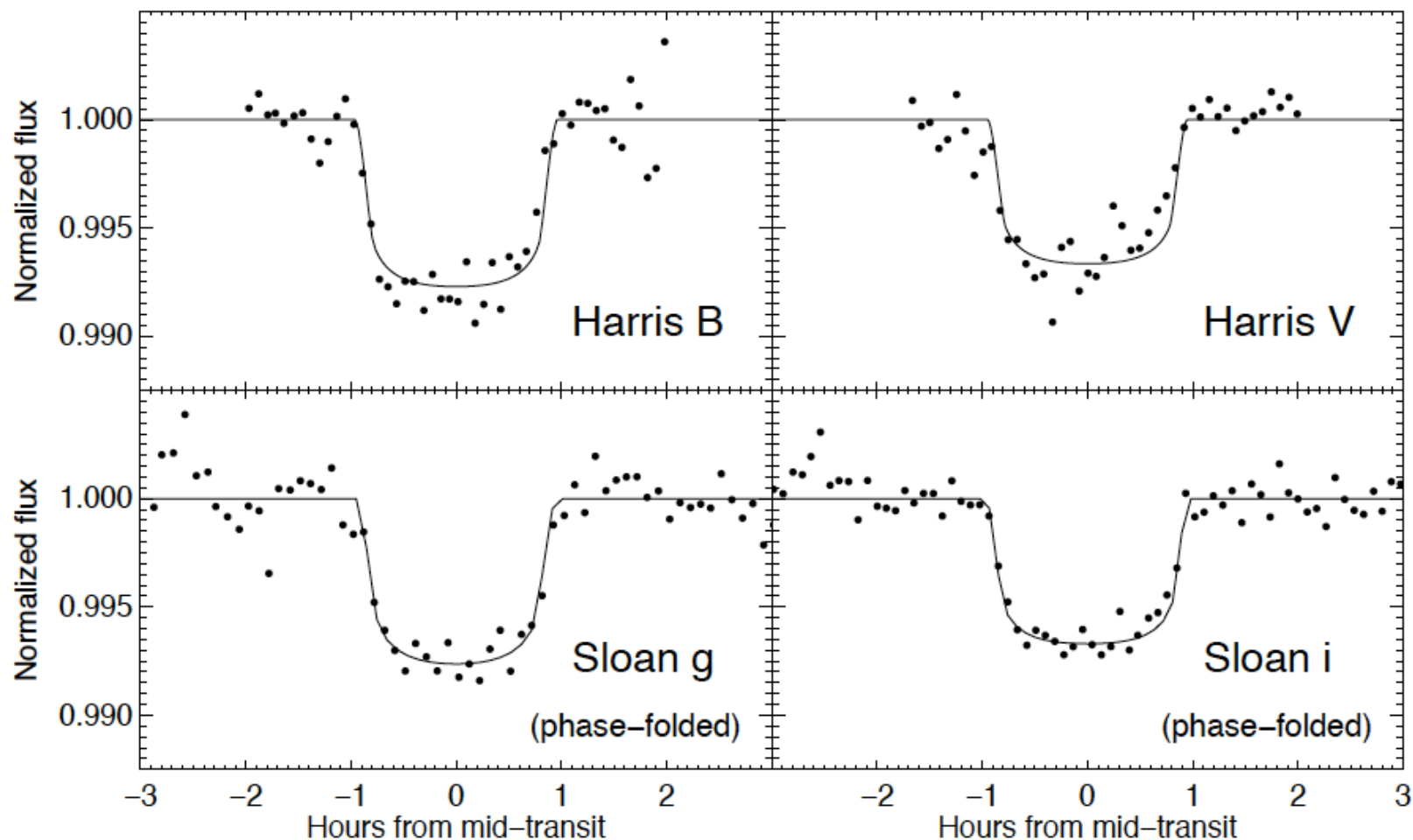
HAT-P 10b transit

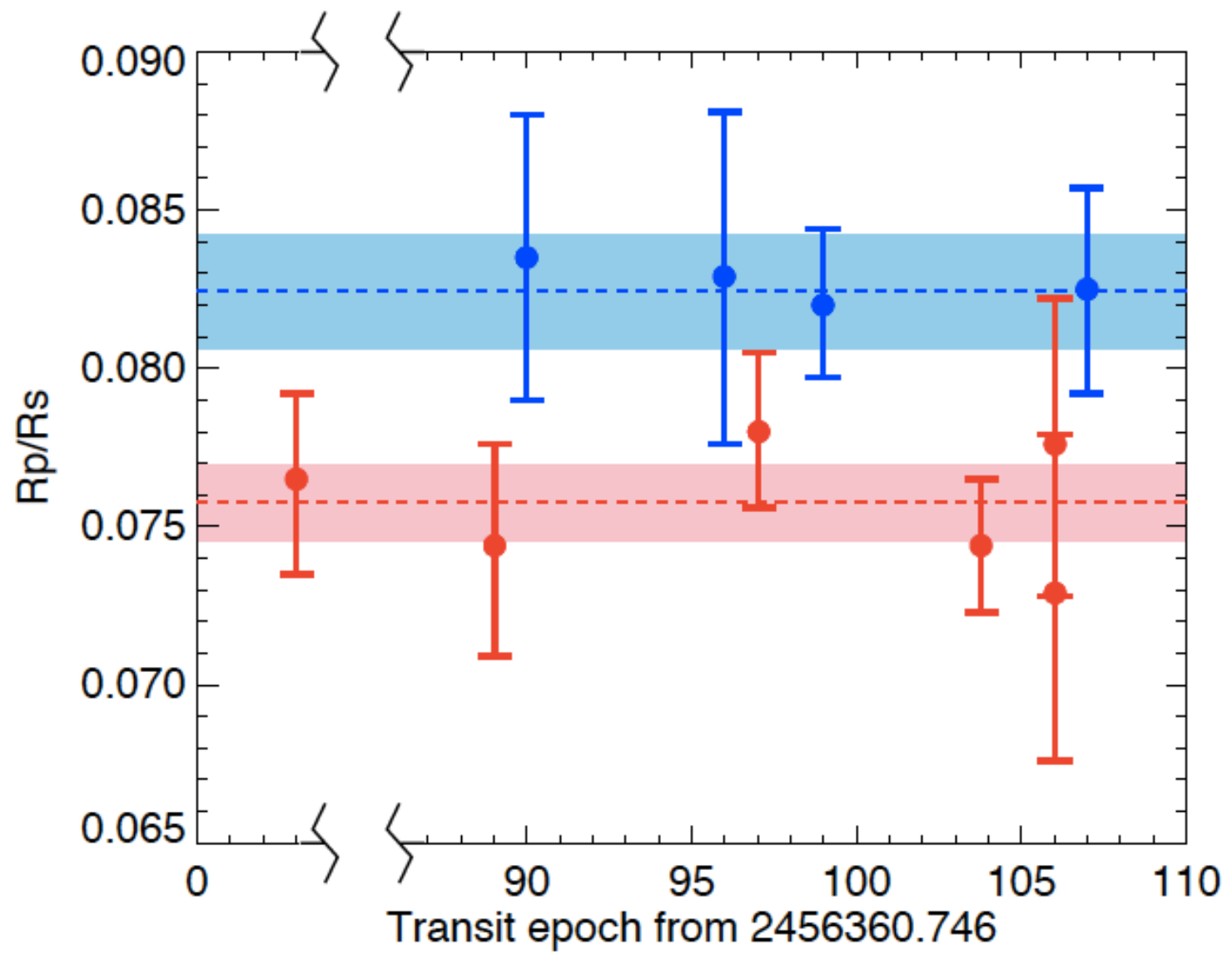


RAYLEIGH SCATTERING IN THE ATMOSPHERE OF THE WARM EXO-NEPTUNE GJ 3470B

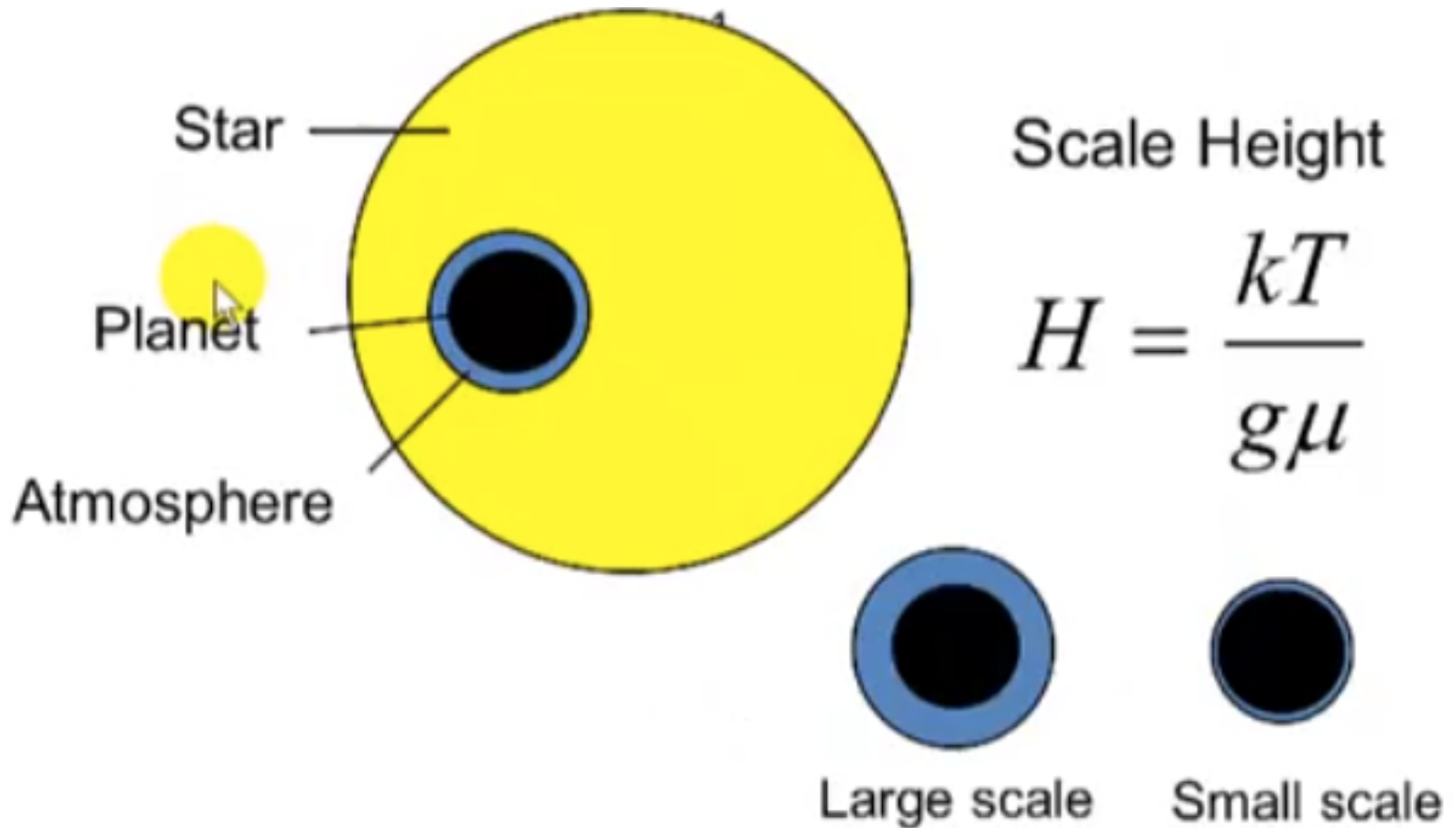
DIANA DRAGOMIR^{1,2,3}, BJÖRN BENNEKE⁴, KYLE A. PEARSON⁵, IAN J. M. CROSSFIELD^{6,7}, JASON EASTMAN⁸, TRAVIS BARMAN⁶, LAUREN I. BIDDLE⁹

Draft version November 19, 2015



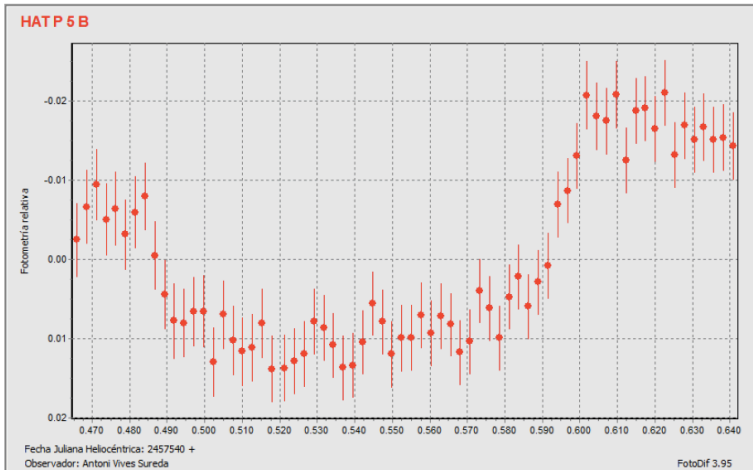


Planetary atmospheres

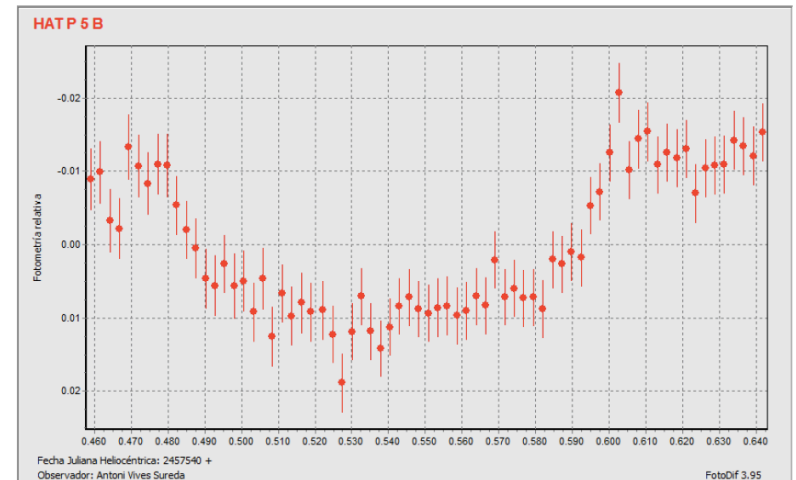


Hat-P-5b, OSN 31 May 2016

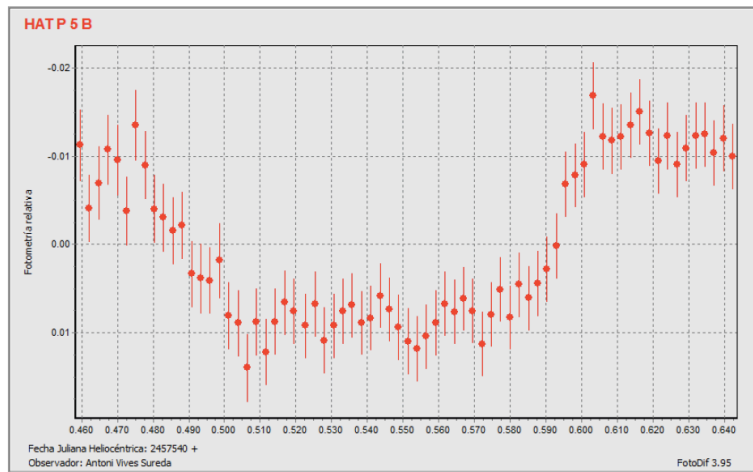
HAT-P-5 b filtro B:



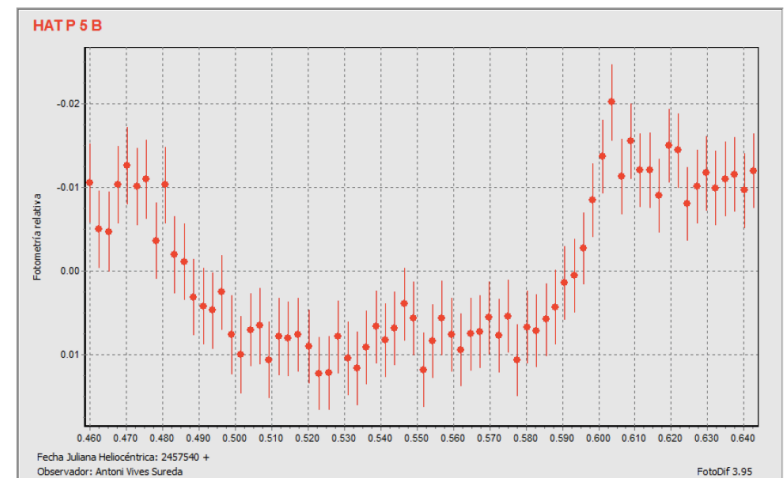
HAT-P-5 b filtro V:



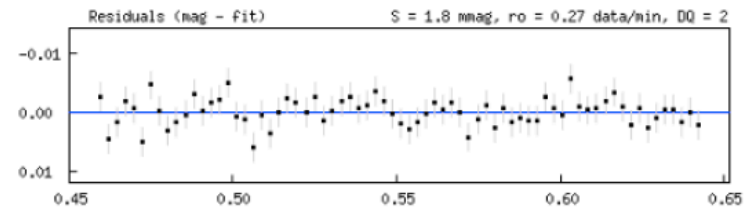
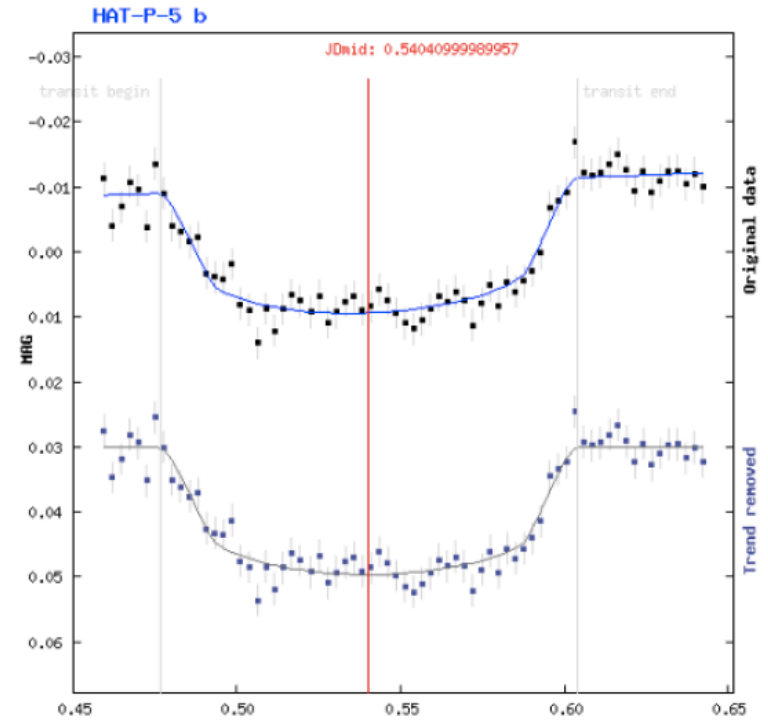
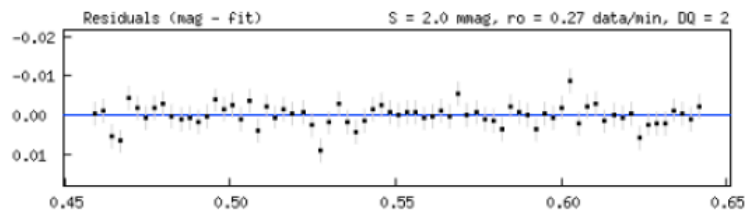
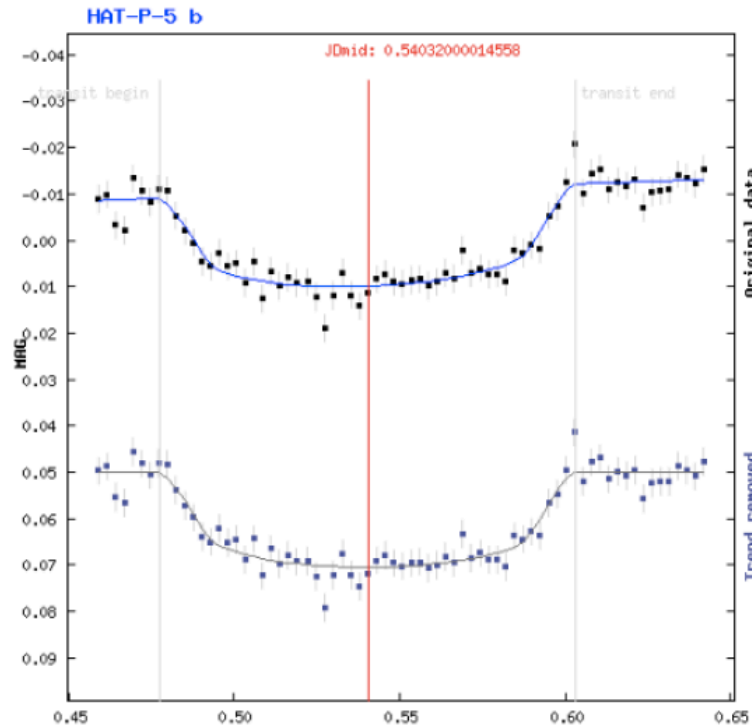
HAT-P-5 b filtro R:



HAT-P-5 b filtro I:



Hat-P-5b, OSN 31 May 2016



Hat-P-5b, OSN 31 May 2016

HAT – P – 5b filtro B:

JD MID: 2457540.53883 \pm 0.00084

HJD MID: 2457540.54147 \pm 0.00084

Duración: 179.6 \pm 2.9 min

Transit depth: 0.0221 \pm 0.0008 mag

HAT – P – 5b filtro R:

JD MID: 2457540.53777 \pm 0.00081

HJD MID: 2457540.54041 \pm 0.00081

Duración: 182.9 \pm 2.8 min

Transit depth: 0.0196 \pm 0.0007 mag

HAT – P – 5b filtro I:

JD MID: 2457540.53725 \pm 0.00089

HJD MID: 2457540.53989 \pm 0.00089

Duración: 185.4 \pm 3.0 min

Transit depth: 0.0198 \pm 0.0008 mag

HAT – P – 5b filtro V:

JD MID: 2457540.53768 \pm 0.00092

HJD MID: 2457540.54032 \pm 0.00092

Duración: 181.2 \pm 3.2 min

Transit depth: 0.0205 \pm 0.0008 mag

WASP-109b, OSN 30 May 2016

FILTRO B

JD mid: 2457539.44233 +/- 0.00120
HJD mid: 2457539.44794 +/- 0.00120 (helcor = **0.00561**)
Duration: 157.4 +/- 4.7 minutes
Depth: 0.0119 +/- 0.0007 mag

FILTRO I

JD mid: 2457539.43992 +/- 0.00132
HJD mid: 2457539.44553 +/- 0.00132 (helcor = **0.00561**)
Duration: 185.0 +/- 5.3 minutes
Depth: 0.0135 +/- 0.0008 mag

FILTRO R

JD mid:	2457539.44347 +/- 0.00485	
HJD mid:	2457539.44908 +/- 0.00485	(helcor = 0.00561)
Duration:	179.4 +/- 16.6	minutes
Depth:	0.0125 +/- 0.0029	mag

FILTRO V

JD mid:	2457539.44083 +/- 0.00227	
HJD mid:	2457539.44644 +/- 0.00227	(helcor = 0.00561)
Duration:	177.0 +/- 7.9	minutes
Depth:	0.0126 +/- 0.0013	mag

TeES-3b, OAO 25 August 2016

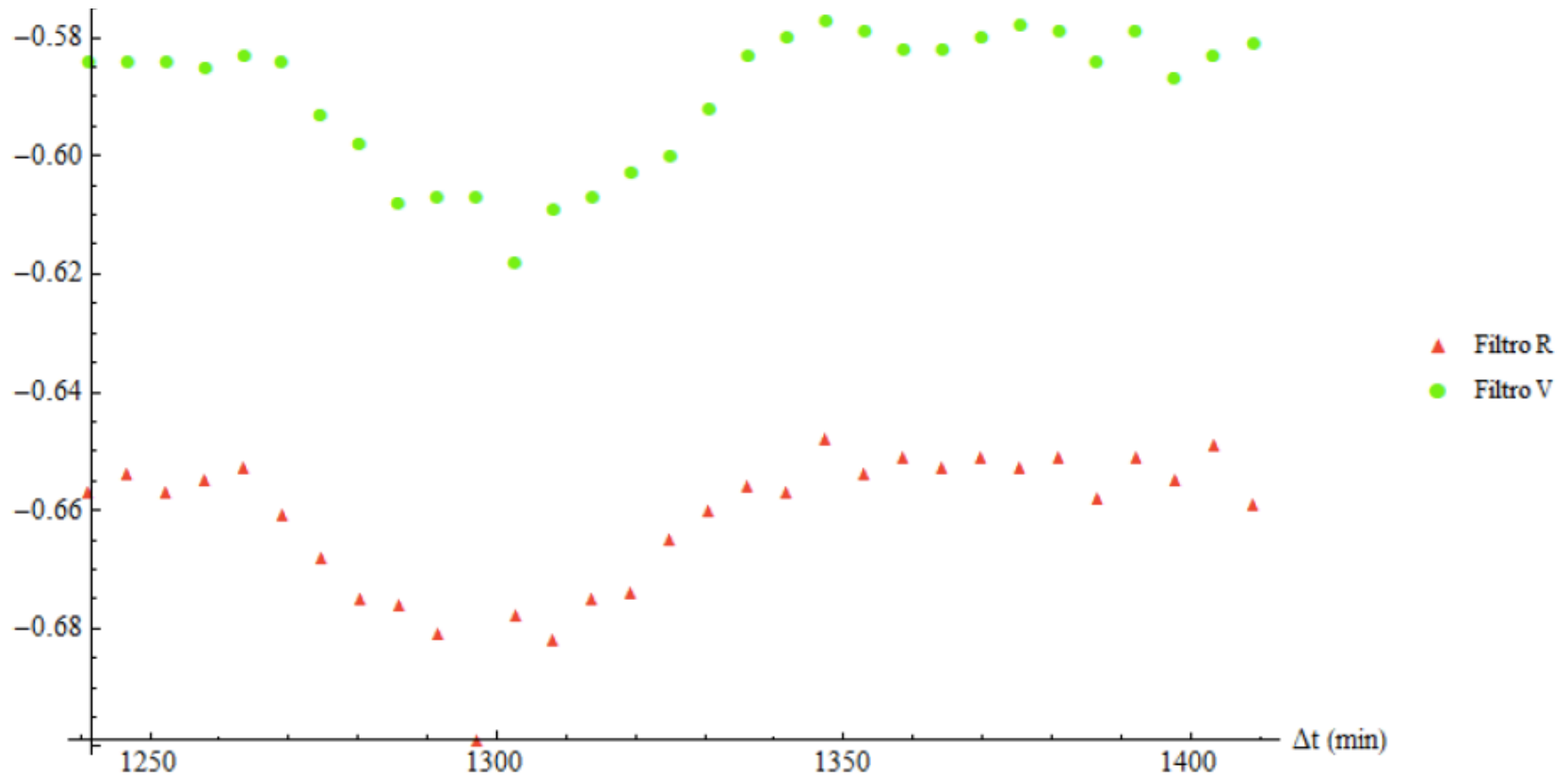


Figura 2: Fotometría diferencial en el filtro R y filtro V.

TeES-3b, OAO 25 August 2016

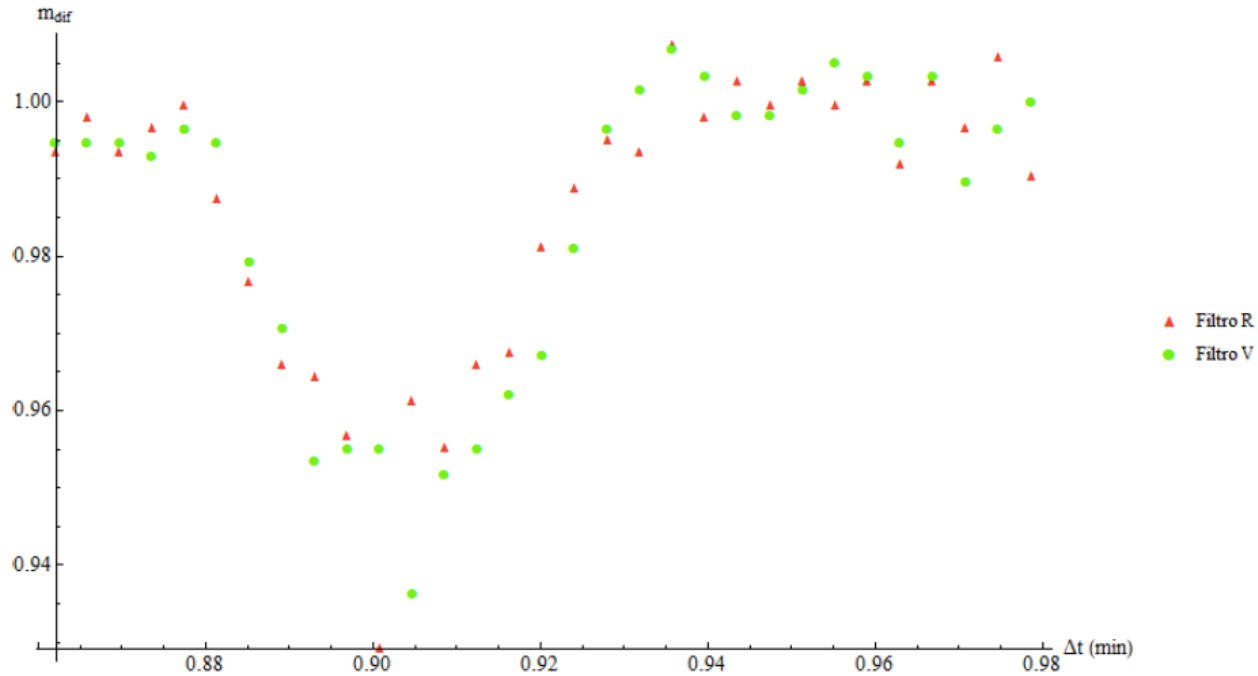


Figura 3: Fotometría diferencial en el filtro R y filtro V normalizada. Se observa una pequeña diferencia en la profundidad del tránsito.

TeES-3b, OAO 25 August 2016

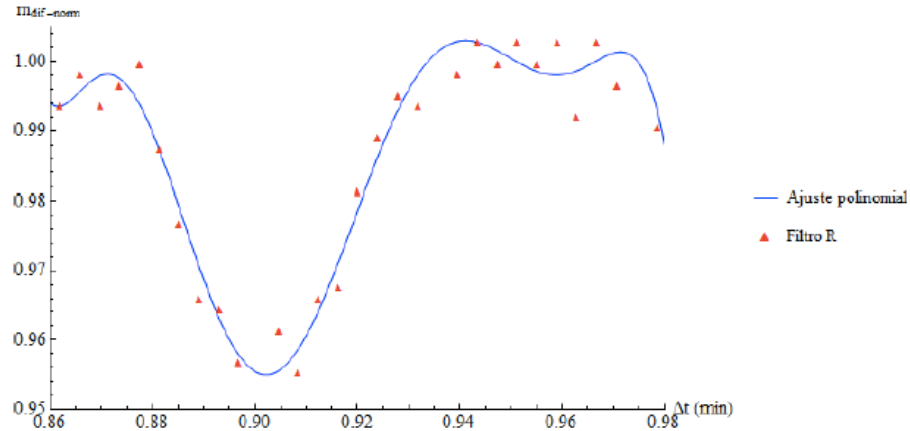
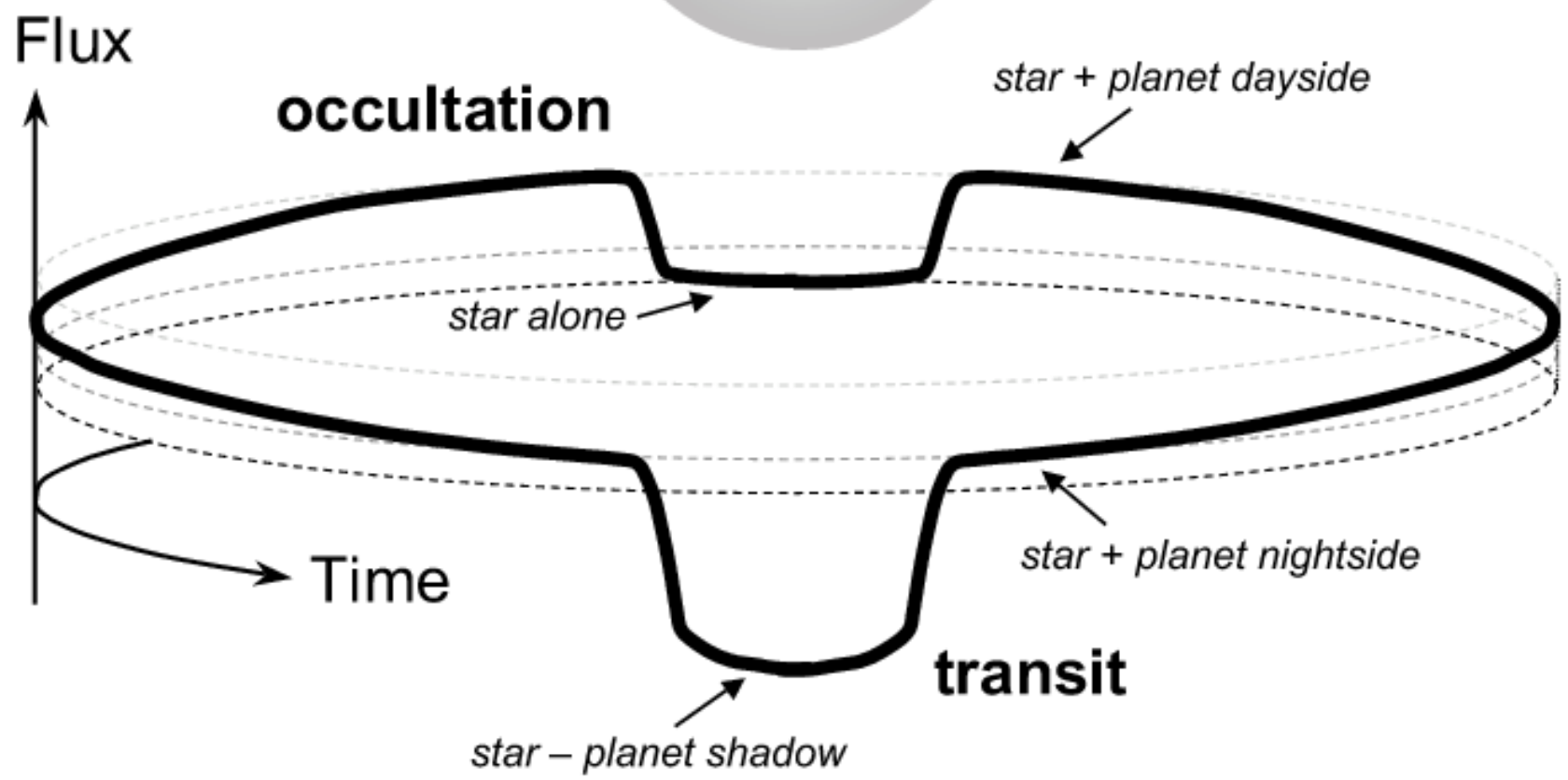
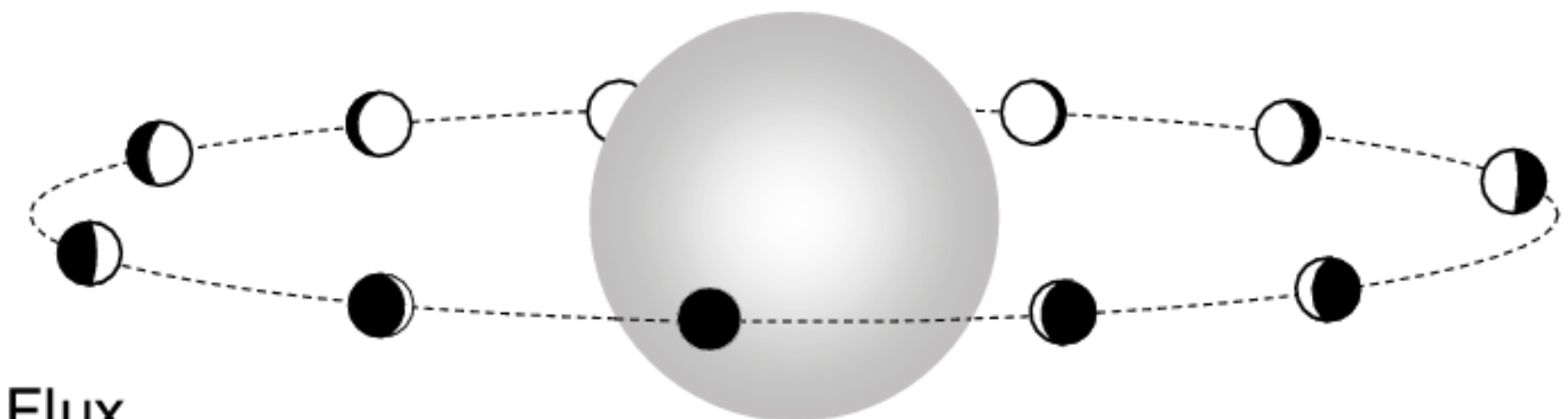
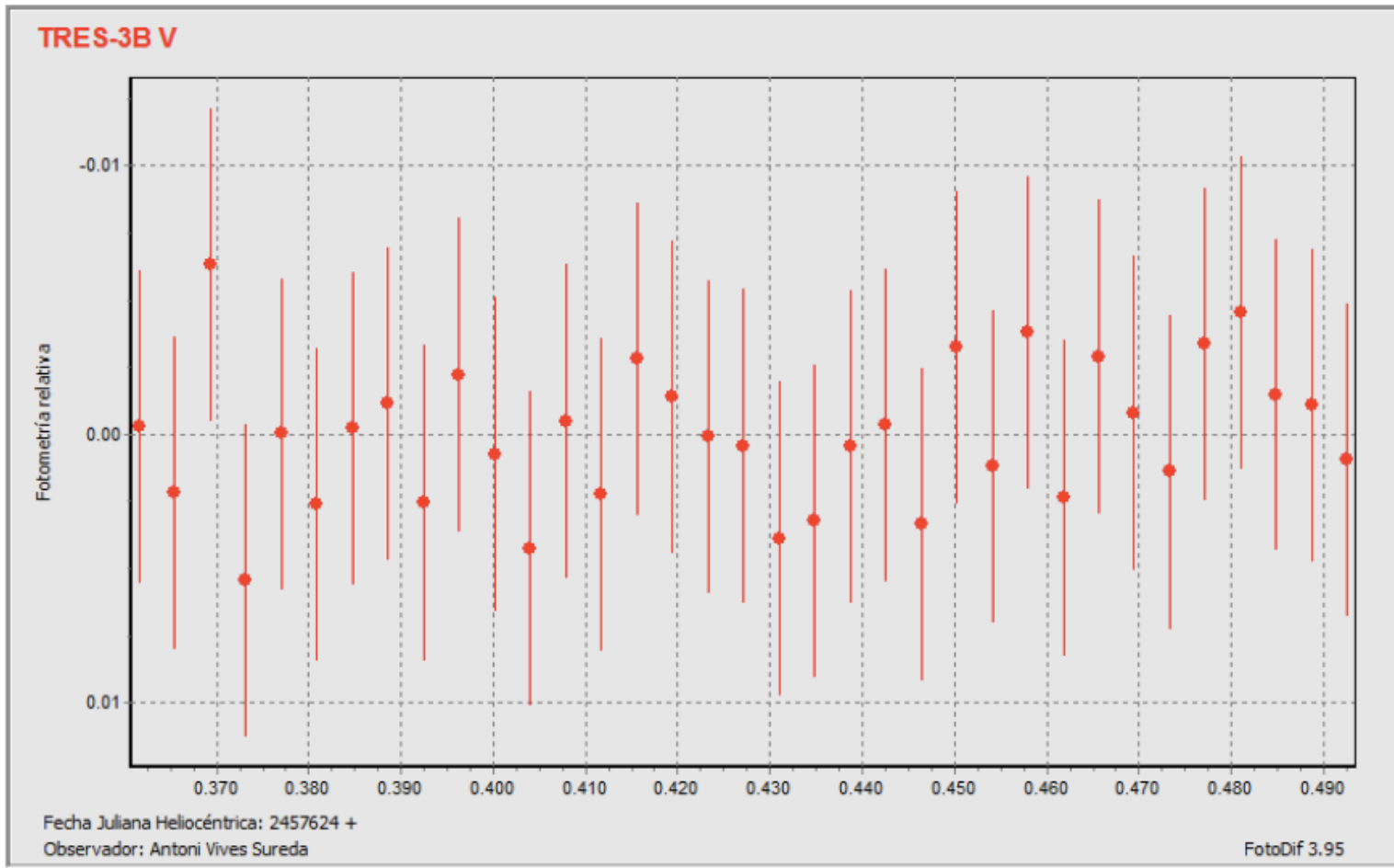


Figura 5: Ajuste polinómico de grado 10 a los datos del filtro R. Hemos suprimido un punto conflictivo. Aumentar el grado del polinomio no proporciona una mejora apreciable.

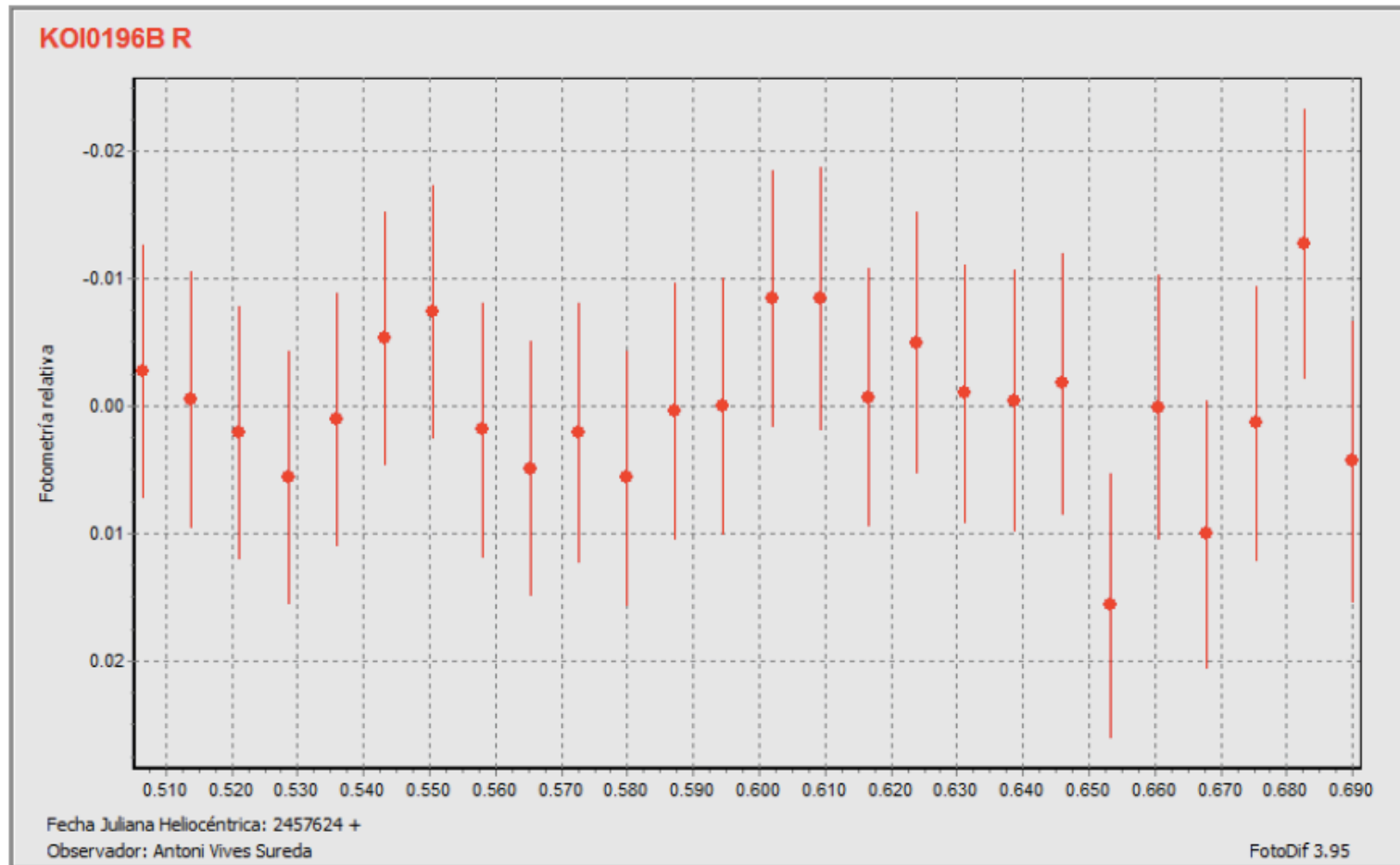
A partir de este ajuste y siguiendo los pasos comentados anteriormente, obtenemos los siguientes valores: $A_V = 0,00172$ y $A_R = 0,00151$, es decir, el área del tránsito en el filtro V es aproximadamente un 10% mayor.



TeES-3b, OAO 23 August 2016



KOI0196b, OAO 25 August 2016



Kepler transits and occultations

Exoplanets	$R_p (R_J)$	i (degrees)	$a(\text{Au})$	Occultation Depth(ppm)	Transit Depth(mag)	Geometric Albedo
Kepler - 5b	1.426 ^{+0.036} -0.051	89.14 ^{+0.44} -0.32	0.0538 ^{+0.0015} -0.0021	18.6 ^{+5.1} -5.3	0.0078 (ETD)	0.121
Kepler - 6b	1.304 ^{+0.018} -0.033	88.93 ^{+0.19} -0.17	0.04852 ^{+0.00074} -0.00133	11.1 ^{+4.8} -5.3	0.0113 (ETD)	0.07
Kepler -7b	1.622 ^{+0.013} -0.013	85.161 ^{+0.055} -0.054	0.06067 ^{+0.00059} -0.0059	39 ⁺¹¹ -11	0.0081 (ETD)	0.248
Kepler- 8b	1.416 ^{+0.053} -0.062	83.978 ^{+0.036} -0.032	0.0474 ^{+0.0018} -0.0021	26 ⁺¹⁰ -10	0.0097 (ETD)	0.133
Kepler- 10b	0.1321 ^{+0.0044} -0.0026	85.1 ^{+3.4} -4.3	0.01720 ^{+0.00081} -0.00168	7.5 ^{+2.0} -2.1	0.000191 (Holczer, 2016)	0.58
Kepler - 12b	1.754 ^{+0.031} -0.036	88.796 ^{+0.088} -0.074	0.0553 ^{+0.0010} -0.0012	20.2 ^{+8.3} -7.6	0.0174 (ETD)	0.092
Kepler - 41b	1.040 ^{+0.035} -0.035	82.214 ^{+0.090} -0.085	0.02396 ^{+0.00081} -0.00080	44 ⁺¹⁵ -16	0.011048 (Holczer, 2016)	0.108
Kepler - 43b	1.219 ^{+0.065} -0.064	84.646 ^{+0.107} -0.091	0.0460 ^{+0.0026} -0.0025	11 ⁺⁴⁰ -37	-----	0.071
Kepler - 76b	1.36 ^{+0.12} -0.12	77.55 ^{+0.20} -0.17	0.0274 ^{+0.0020} -0.0019	131.6 ^{+8.7} -8.0	-----	0.246
Kepler - 91b	1.367 ^{+0.069} -0.060	69.68 ^{+0.67} -0.57	0.0731 ^{+0.0034} -0.0031	35 ⁺¹⁸ -18	0.000472	0.46
Kepler - 412b	1.341 ^{+0.044} -0.046	80.559 ^{+0.084} -0.079	0.02897 ^{+0.00093} -0.00092	53 ⁺²¹ -24	0.010533 (Holczer, 2016)	0.113
TrES - 2b	1.247 ^{+0.050} -0.045	83.872 ^{+0.020} -0.018	0.0367 ^{+0.0014} -0.0013	7.7 ^{+2.4} -2.6	0.0169	0.031
HAT - p7b	1.419 ^{+0.178} -0.085	83.143 ^{+0.023} -0.02	0.0355 ^{+0.0045} -0.0021	71.2 ^{+1.9} -2.2	0.0067	0.2044
KOI - 13b o Kepler - 13 A b	1.512 ^{+0.035} -0.035	86.770 ^{+0.048} -0.052	0.03641 ^{+0.00087} -0.00087	172.0 ^{+1.7} -1.6	0.0046	0.4565
KOI - 13b o Kepler - 13 A b	2.63 ^{+1.04} -0.82	86.769 ^{+0.046} -0.050	0.063 ^{+0.0025} -0.0020	170.8 ^{+1.6} -1.5	0.0046	0.4532
KOI - 13b o Kepler - 13 A b	2.216 ^{+0.087} -0.087	86.770 ^{+0.047} -0.046	0.0533 ^{+0.0021} -0.0022	170.8 ^{+1.6} -1.5	0.0046	0.4529

Conclusions

- Differences in transit curve depths observed
- A photometric accuracy of ~ 0.003 mag or better needed
- Results to be confirmed. Long term programme.